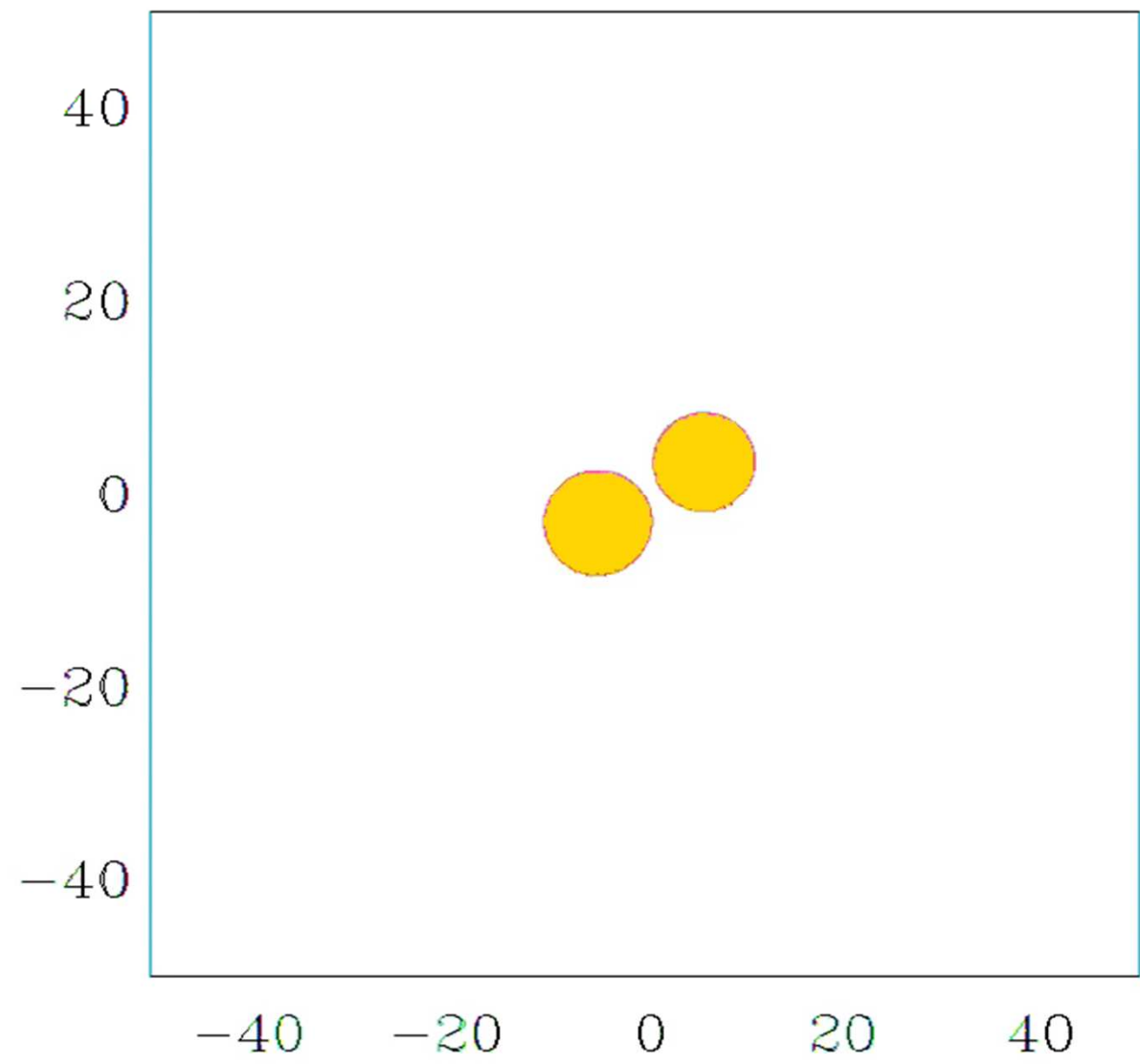


The year is 100,000,000. (The age of the Earth and of the solar system, is $\sim 4,568,000,000$)



SwRI

Canup (2012)

ABINIT and the formation of the Moon

Razvan Caracas
CNRS, ENS de Lyon

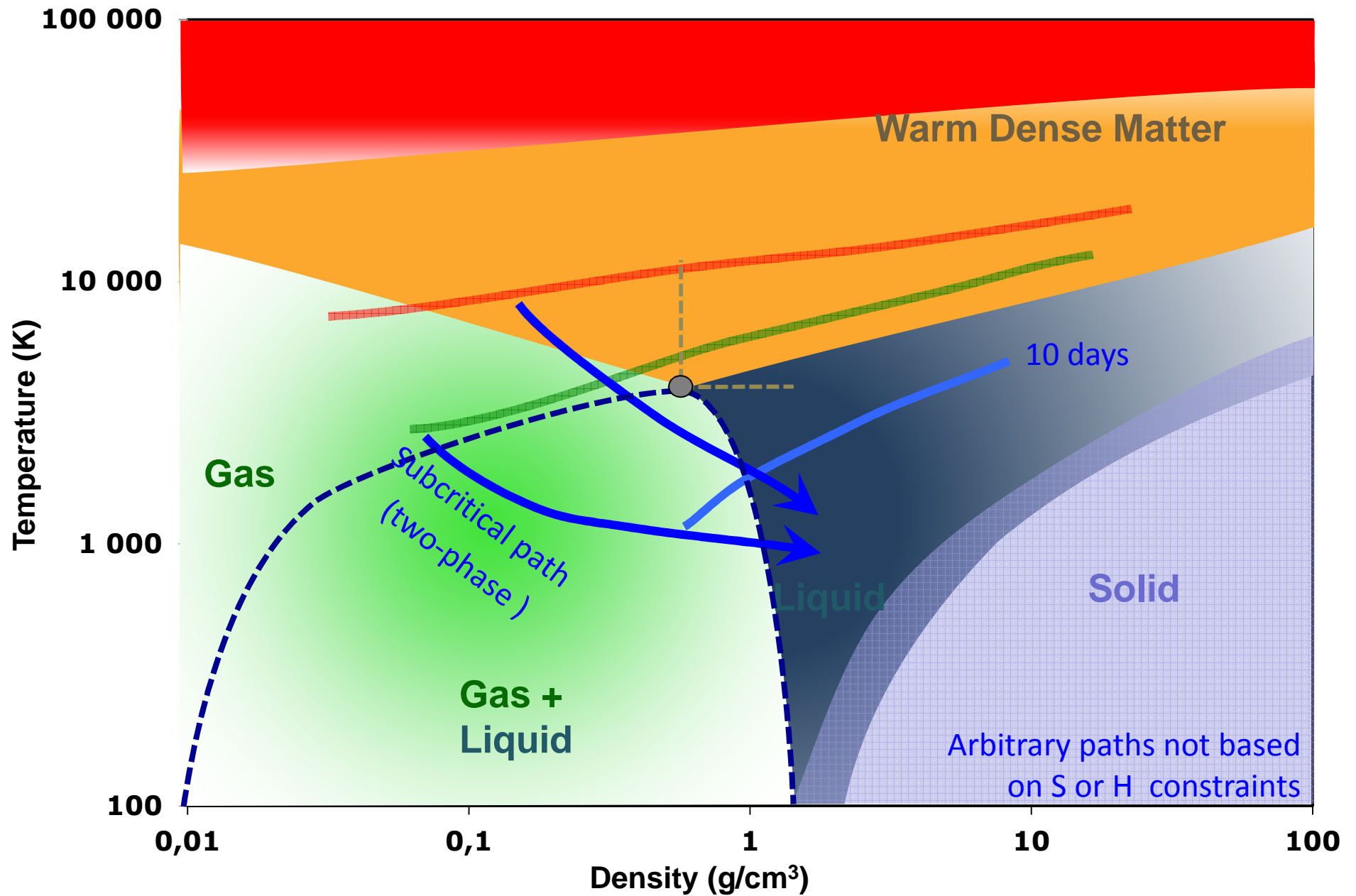


The logo for the IMPACT project, consisting of a small grey sphere with a white crater-like feature to the left of the word 'IMPACT' in a large, bold, red, sans-serif font.



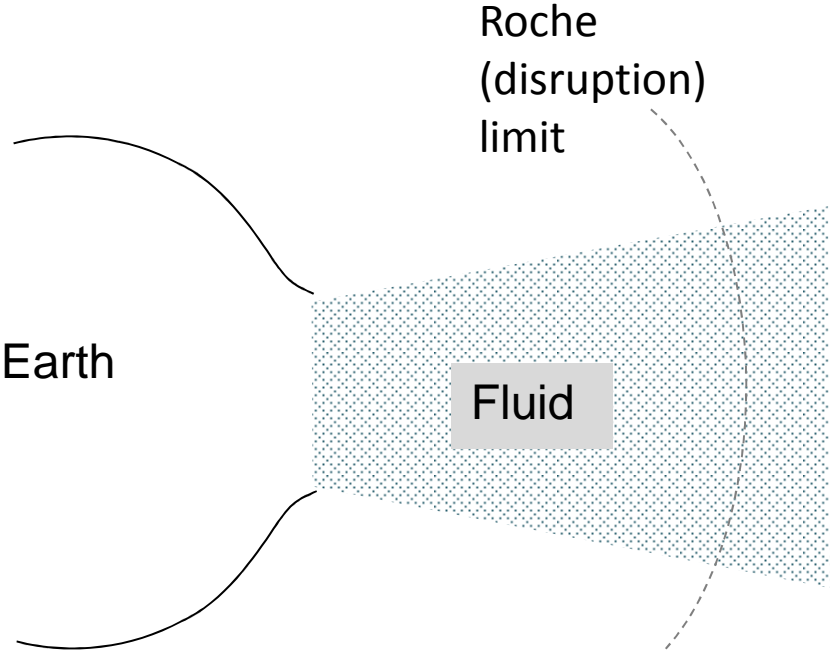
*Evolution and condensation
of the proto-lunar disk*

subcritical or supercritical? this is the question!

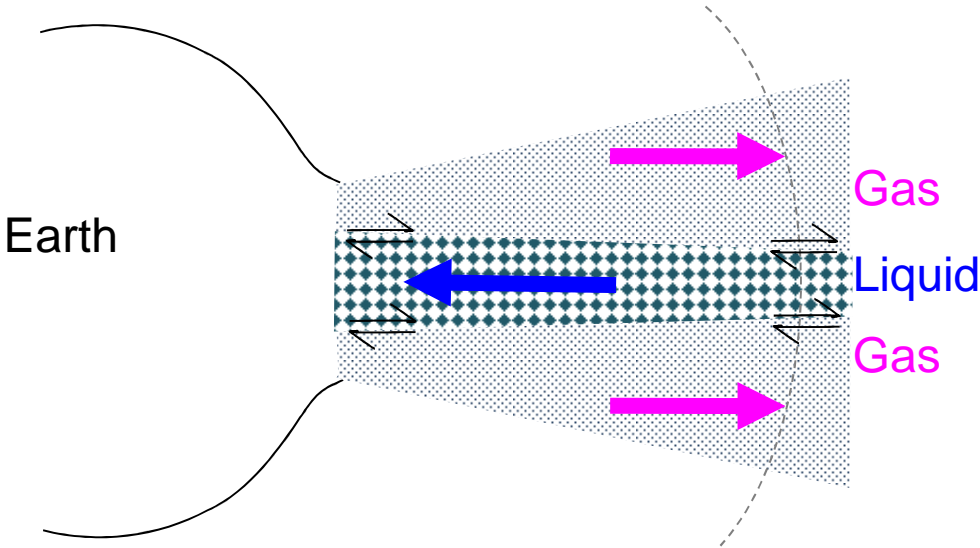


Different dynamical regimes of the disk after the impact

*Exchange of angular momentum
by liquid – vapor friction*

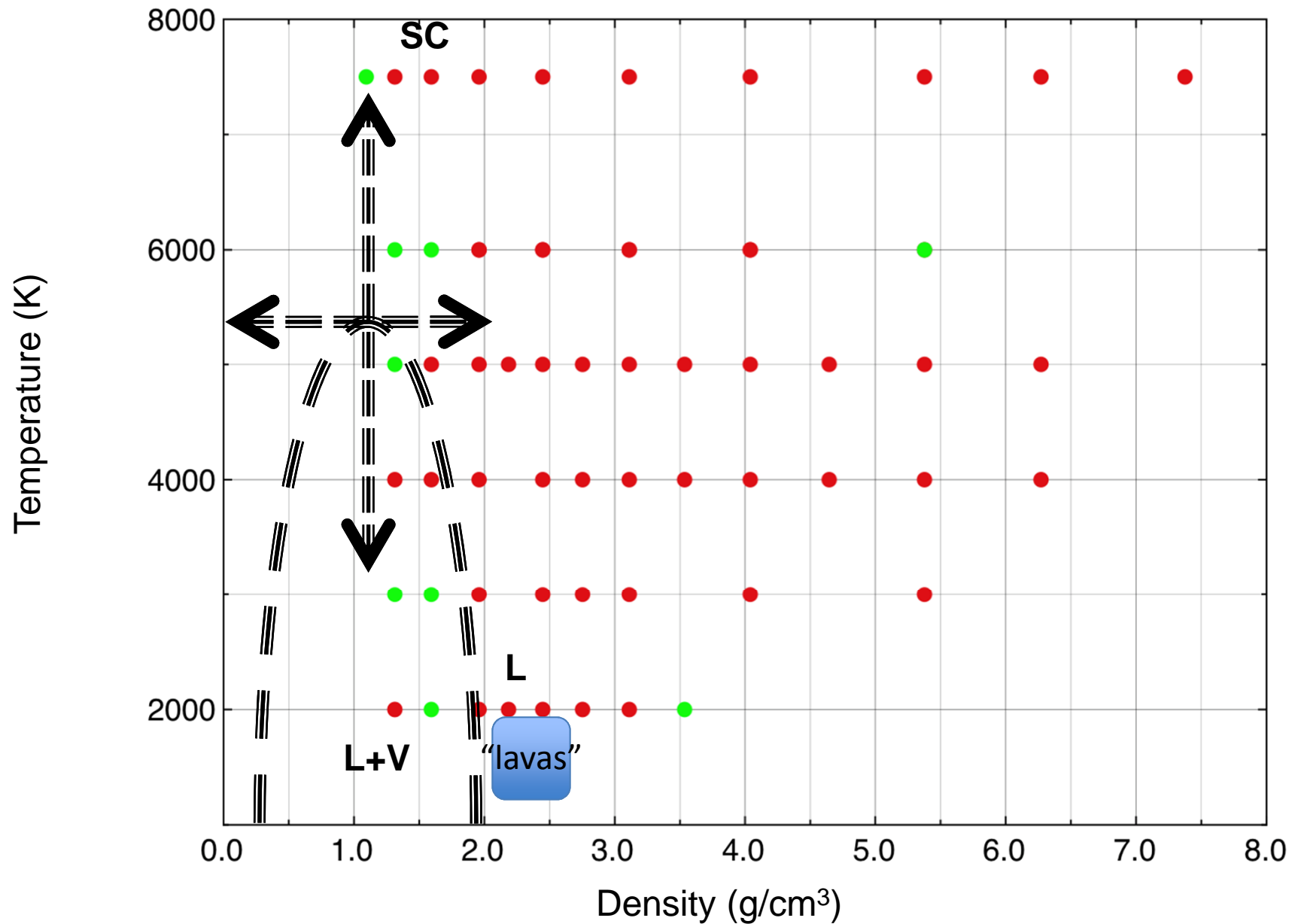


HOTTER disk
Supercritical
one-phase

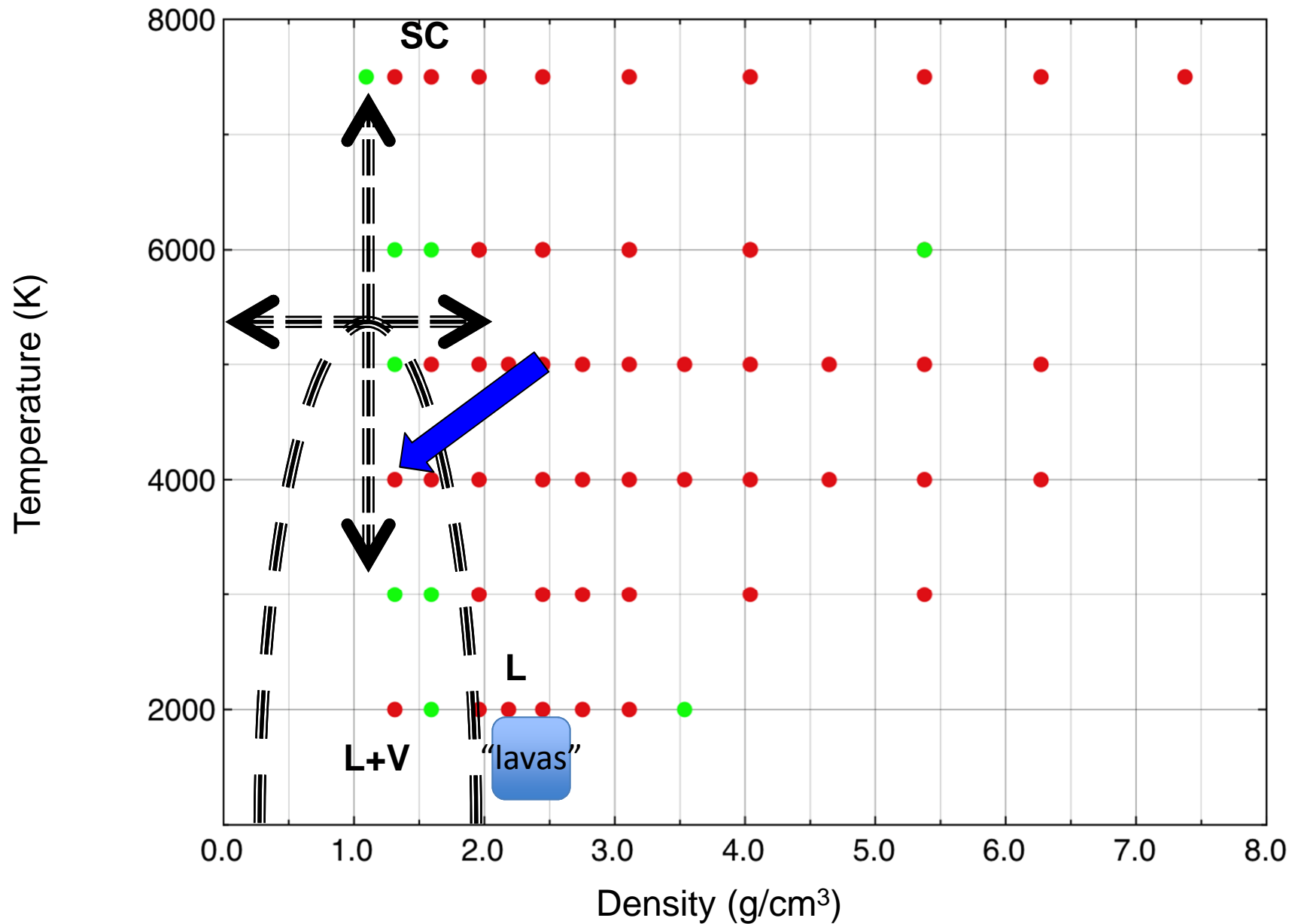


COLDER disk
Subcritical
two-phase

RECONSTRUCTING THE GIANT IMPACT - pyrolite melts

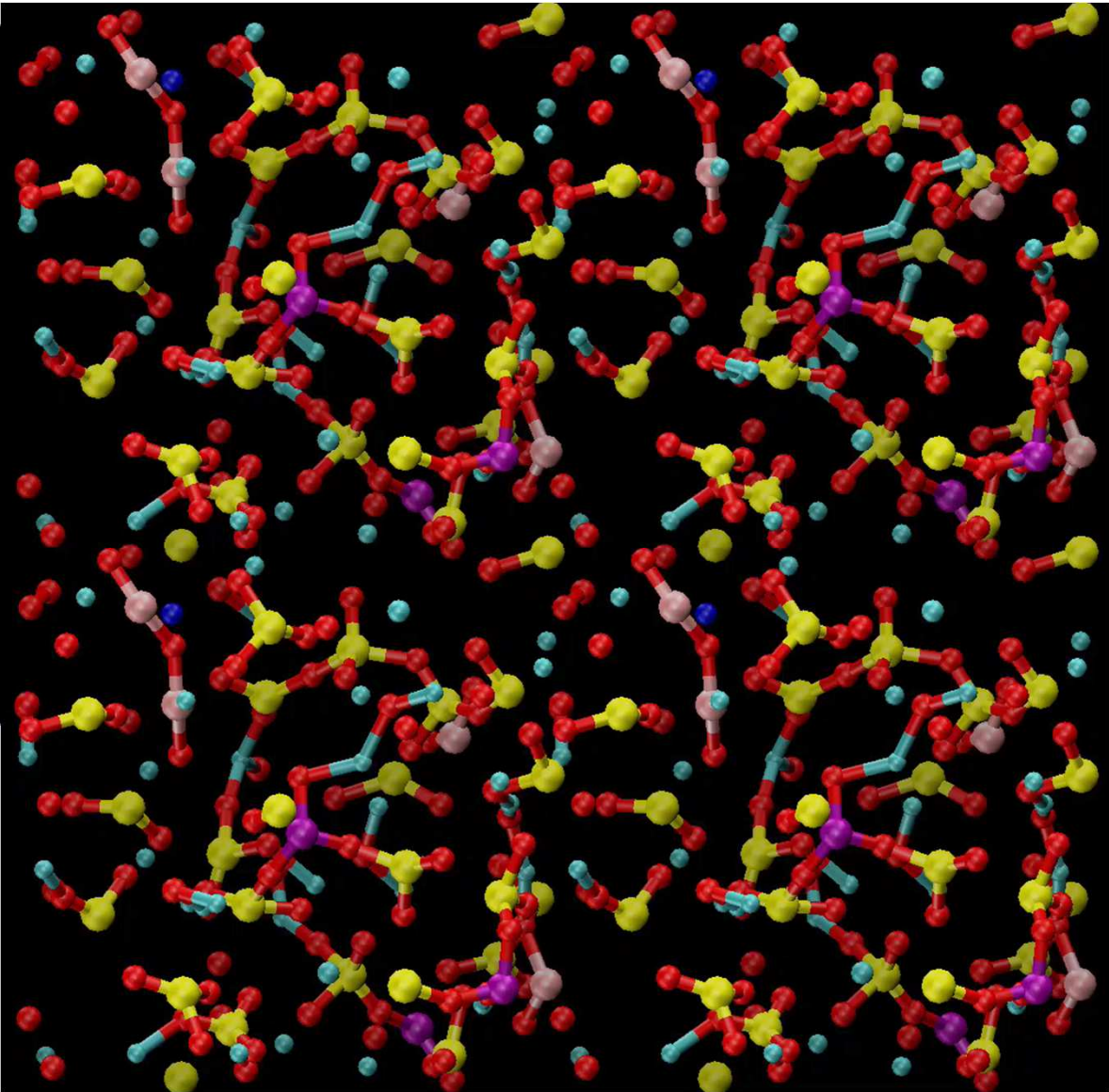


RECONSTRUCTING THE GIANT IMPACT - pyrolite melts

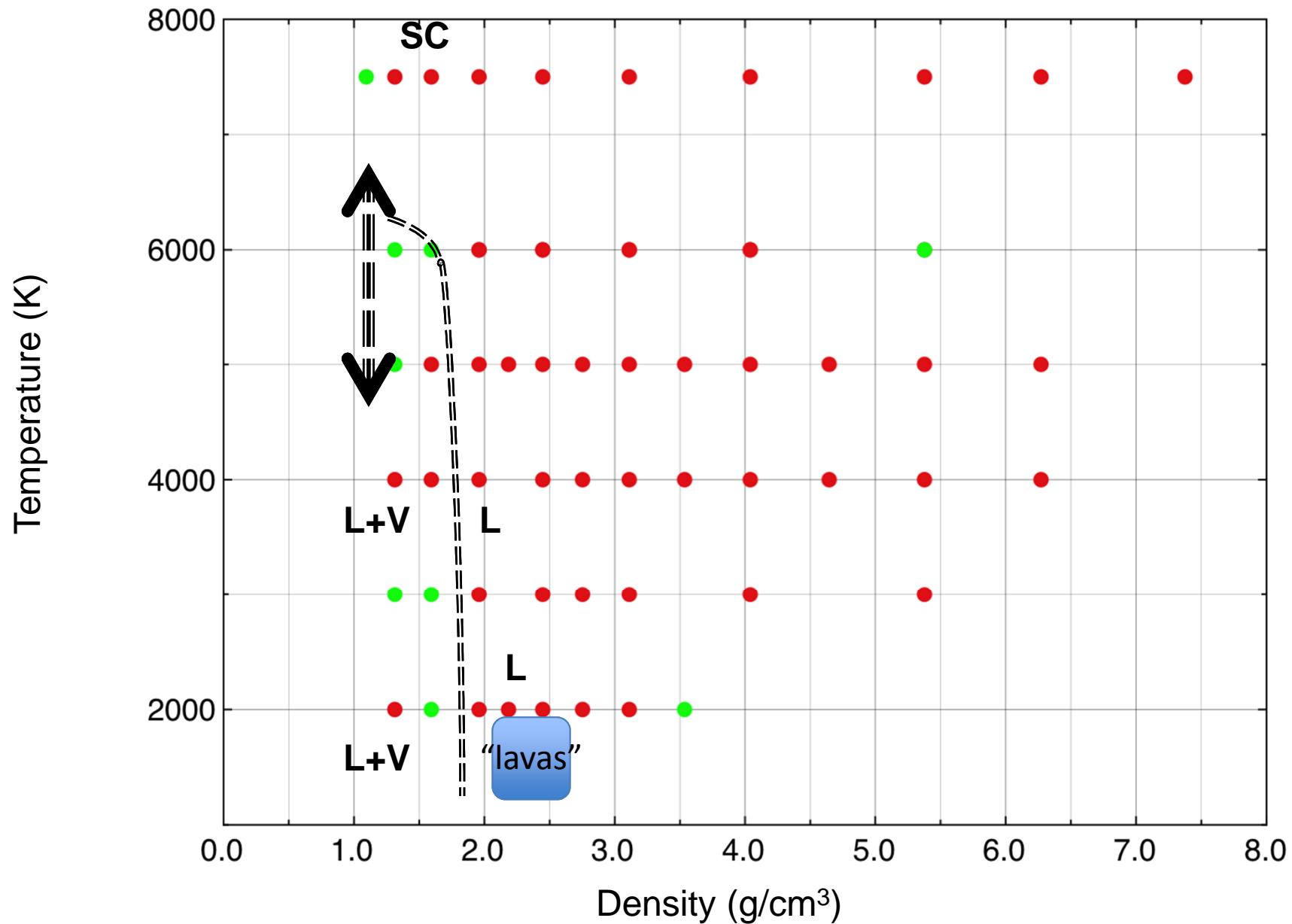


$\rho = 1.313 \text{ g/cm}^3$
 $T = 4000\text{K}$

Mg
Fe
Ca
Na
Al
Si
O



RECONSTRUCTING THE GIANT IMPACT - pyrolite melts



TO DO list:

1. (Thermo)dynamics from (PI)MD simulations
 - Compute the supercritical domes
 - Place the supercritical points
 - Build the liquid/vapor equilibrium “curves”
 - Equation of states of the liquid
 - Thermodynamic properties of the liquid and gas
 - Speciation in the vapor phase
 - Isotope and chemical partitioning functions
2. Electronic/optical properties
 - GW calculations on atomic clusters of the gas phase
 - GW calculations on the liquid phase (snapshots)
 - Optical signatures for identification